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#### **PAPER**

# Diagnosis of Weibel instability evolution in the rear surface density scale lengths of laser solid interactions via proton acceleration

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#### Abstract

It is shown for the first time that the spatial and temporal distribution of laser accelerated protons can be used as a diagnostic of Weibel instability presence and evolution in the rear surface scale lengths of a solid density target. Numerical modelling shows that when a fast electron beam is injected into a decreasing density gradient on the target rear side, a magnetic instability is seeded with an evolution which is strongly dependent on the density scale length. This is manifested in the acceleration of a filamented proton beam, where the degree of filamentation is also found to be dependent on the target rear scale length. Furthermore, the energy dependent spatial distribution of the accelerated proton beam is shown to provide information on the instability evolution on the picosecond timescale over which the protons are accelerated. Experimentally, this is investigated by using a controlled prepulse to introduce a target rear scale length, which is varied by altering the time delay with respect to the main pulse, and similar trends are measured. This work is particularly pertinent to applications using laser pulse durations of tens of picoseconds, or where a micron level density scale length is present on the rear of a solid target, such as proton-driven fast ignition, as the resultant instability may affect the uniformity of fuel energy coupling.

# 1. Introduction

A plethora of laser accelerated proton beam applications have been the subject of intense speculation since the target normal sheath acceleration mechanism heralded beams with unrivalled emittance, source bunch lengths and densities.

These highly laminar beams often have a uniform spatial distribution which is essential for picosecond probing of intense transient fields [1], and their high degree of focusability lends them to potential applications such as hadron therapy [2, 3] and fast ignition [4], where the high currents of laser accelerated beams are uniquely capable of delivering the power densities required for the latter.

In theorised fast ignition scenarios, kilojoule level chirped pulse amplification systems are often envisaged with pulse durations of tens of picoseconds [4] to accelerate an ignitor ion beam. We report on experimental data where a double pulse interaction opens a new route to diagnosing instabilities in the acceleration dynamics that may be encountered on such systems, where pulse delays on the order of tens of picoseconds are introduced<sup>4</sup>. We observe that the laminarity of proton acceleration in this regime is reduced, through Weibel instability growth in the target rear scale lengths. This is diagnosed by a filamentation of the proton beam accelerated during interactions with 25  $\mu$ m gold and 45  $\mu$ m graphite foils.

 $<sup>^4</sup>$  We note here, where the beams have an intensity ratio of 1:10, the temporal separation between them acts as a sample of a full width at tenth maximum (FWTM) of a Gaussian distribution, this is longer than the often quoted, full width at half maximum (FWHM), single pulse duration. The two values are related by FWTM =  $(\ln(10)/\ln(2))^{\frac{1}{2}}$  FWHM or FWTM  $\approx 1.8$  FWHM. The quoted time delay can therefore be considered to be a sample of the intensity profile of a shorter duration FWHM pulse.

Studies have consistently shown that the distribution of the laser accelerated proton beam is dependent on that of the laser driven hot electron population, and many mechanisms have been identified as being responsible for the electron beam itself gaining spatial structure.

Extensive numerical modelling and experimental studies have led to a powerfully predictive 3D hybrid model showing that low temperature plasma (10–100 eV) resistivity gradients can lead to filamentation of the electron beam [5–7]. Using the final electron distribution at the target rear, a sheath expansion model [8] has been shown to accurately forecast the subsequently accelerated proton distribution, and this work established a precedent for using the accelerated proton beam spatial distribution as a diagnostic of fast electron transport physics within the target.

The underlying physics of this mechanism requires that targets of thickness in excess of  $100 \mu m$ , are required for the necessary longitudinal temperature gradients to exist for electrically non-conductive materials to enter a resistivity regime deviating from Spitzer or Lee-More [6]. To this end, the model predicts this mechanism should have little effect on metallic foils, and experimentally no attributable ion beam structure in metallic foils of up to 1 mm thick has been observed [9].

Metzkes  $et\,al\,[10]$  have reported on structured proton beams, using  $1-2\,\mu\mathrm{m}$  titanium foils, where a small preplasma, optimised for scale lengths of  $0.1\,\mu\mathrm{m}$ , on the target front was attributed to being the source of a Weibel instability similar in nature to that reported in a numerical study by Sentoku and Kemp [11]. Experimentally [10] a target thickness of  $3-5\,\mu\mathrm{m}$ , was established as a limit on the thickness through which this front surface seeding of electron beam structure would be transported to the target rear, as the proton beam is observed to return to a smooth spatial distribution.

In a series of experimental investigations into longer front surface preplasma scale lengths, ranging from 10s to 100s  $\mu$ m [12–14], no evidence has arisen to suggest that any front surface seeding of electron beam filamentation is transported through 5–25  $\mu$ m copper targets. Here the proton beams were shown to be featureless and smoothest for the largest scale length preplasma conditions. In a closely related numerical investigation of these experimental parameters [15, 16] it was found that at the largest scale lengths, the incident laser filaments in the preplasma. This explains the observed reduction in maximum proton energy, but is observed to have no effect on the proton beam structure, even when it has been shown that electron beams can attain structre due to front surface illumination nonuniformities [17].

Tatarakis *et al* [18] have shown experimentally, that a Weibel instability can result from driving a dense electron population through a low density plasma. We find a similar phenomenon in the rear surface scale lengths of solid targets and measure the effects on the laser accelerated protons for the first time. In the context of the experimental results presented here, the scale length is understood to emanate from the bulk target expansion initiated by the heating of the target foil by the controlled prepulse, and longer scale lengths therefore correspond to longer pulse delays. A detailed discussion of the relation between these parameters will be presented in section 4.2.

In previous studies, the picosecond evolution of the Weibel instability has not been explored, which is an important time frame for many high power laser systems and for ion acceleration. We present numerical results, showing how the evolution varies as a function of scale length on the picosecond time frame, and present experimental evidence that the time history of this evolution is mapped onto the accelerated proton beam.

As well as being an important consideration for ion acceleration using tens of picosecond pulses, the mapping of structure onto the proton beam presents a novel method for diagnosing the onset of the Weibel instability at near solid density, an environment not easily probed by optical methods.

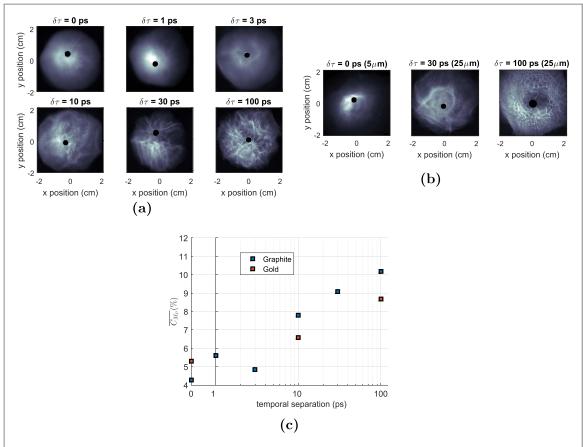
# 2. Experimental details

# 2.1. Methods

The experiment was carried out on the PHELIX laser at the GSI facility [19]. The s-polarised, 1.054  $\mu$ m wavelength, 500 fs FWHM laser pulses were delivered to foil targets with an f/9 off-axis parabola and focused to a 20  $\times$  15  $\mu$ m spot at a 34° angle of incidence. An ultrashort optical parametric amplification module was used in the front end of the laser, offering an intensity contrast ratio of 10° on the nanosecond timescale [20].

A double pulse was generated via the same module described in previous work, which investigated the optimisation of plasma mirror reflectivity [21], delivering a prepulse to main pulse energy of 1:10 and with variable delay of up to 100 ps. In total,  $72\pm2$  J of laser energy was delivered on target for the shots considered here, giving an average focused main pulse intensity of  $4\times10^{18}$  W cm $^{-2}$ . The accelerated proton beam spatial intensity distribution was recorded using  $51\times51$  mm pieces of HDv2 radiochromic film, arranged in a stack configuration for energy discrimination, and positioned 51 mm from the target rear.

Graphite foils of thickness 45  $\mu$ m were used, where a comprehensive scan was obtained, varying the interpulse temporal separation between 0 ps (single pulse) and 100 ps. As well as this, shots were taken on a



**Figure 1.** The dose maps retrieved for 6.5 MeV protons, as measured on the RCF stack, from interactions with (a) 45  $\mu$ m graphite foils (b) 5–25  $\mu$ m gold foils, at a series of time delays between the pulses. (c) The layer averaged mean coefficient of variation as derived from [8] is shown for the two data sets, as a quantitative measure of the observed structure.

combination of gold foils of thickness 5 and 25  $\mu$ m, which is a more sparsely populated data set, but important for our identification of the physical mechanism at work.

#### 2.2. Results

In figures 1(a) and (b), the 2D proton spatio-dose distributions are shown for various inter pulse time delays. Qualitatively, it is clear that the beam changes from a smooth centrally peaked profile, for inter pulse time delays in the region of 0-3 ps to a beam that shows a high degree of filamentation within the envelope for longer time delays.

To quantify the degree of filamentation within the beam, the method used by MacLellan *et al* [8] is employed, where the mean coefficient of variation,  $C_{M_p}$ , represents as a percentage, the radially averaged standard deviation in proton dose divided by the radially averaged dose. This measurement allows the degree of disorder in the beam to be quantified as a function of its radius, and as we find this to be quasi-constant, we take the average value of this function,  $\overline{C_{M_p}}$ , to represent the degree of disorder in the beam measured at a given proton energy. This is shown in figure 1(c) and indicates that the degree of disorder in the beam tends to increase as the inter pulse time delay increases, for both 45  $\mu$ m graphite foils and 25  $\mu$ m gold foils.

By examining the degree of structure in the proton beam as a function of energy, we examine its evolution, as higher energy protons are accelerated at early time when the target rear surface electric field is at its peak and lower energy protons are accelerated later in time as the accelerating field slowly decays. This data is shown in figure 2.

From this plot it may be inferred that the nature of the phenomenon which causes the structure has different time histories for each inter pulse time delay. In the case of 0 ps delay, no substructure is observed in the beam envelope, and the trend in disorder decreasing with higher proton energy informs us on changes in the global spatial distribution at each energy, as a reference case. For the other delays shown, the disorder increases as a result of the substructure within the beam envelope. Comparing the two longer inter pulse time delays, a large degree of substructure is imprinted on the beam at high energy, and maintains this high degree of disorder from early time. The 10 ps case however shows a degree less structure than the 100 ps data. This provides strong

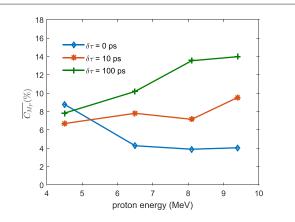


Figure 2. Structure in the beam as a function of proton energy for inter pulse time delays of 0, 10 and 100 ps from interactions with 45  $\mu$ m graphite foils.

evidence that there is an evolutionary difference in the mechanism causing the structure, on the same time scale as the ion acceleration mechanism.

# 3. Numerical modelling

#### 3.1. Two dimensional particle in cell

To investigate the effects of rear surface scale lengths on the interaction dynamics, 2D particle in cell modelling was carried out using the EPOCH code [22]. Here a 500 fs, FWHM Gaussian laser pulse propagates in the positive *x*-direction, centred on the *y*-axis with a spot width of 7  $\mu$ m. The simulation box dimensions were 63(*x*)  $\mu$ m × 24(*y*)  $\mu$ m, with a grid resolution of 40 cells per micron.

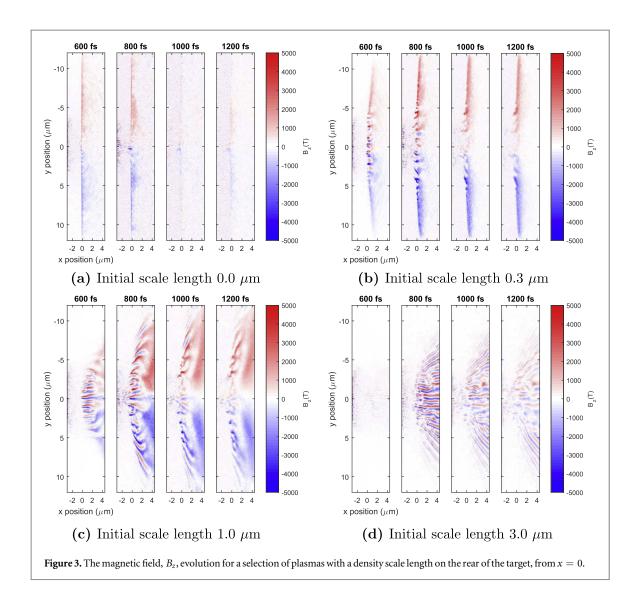
The laser interacts with a target of maximum electron density of 50 times the critical density, spanning a thickness of 3  $\mu$ m in the x direction. The collisionless target is composed of a mixture of ions of charge to mass ratio, 1 and  $\frac{1}{12}$ , representing the contaminant hydrogen layer and bulk target carbon ions, present in a ratio of 1:9 respectively and is quasi-neutralised by electrons with an initial 500 eV temperature. Each species is represented by 100 macroparticles per cell. A scale length of the form  $n_{\rm e;i} = n_{\rm e0;i0} \exp{(-x/l_{\rm ss})}$  is introduced on the target rear surface at  $x=0~\mu$ m.

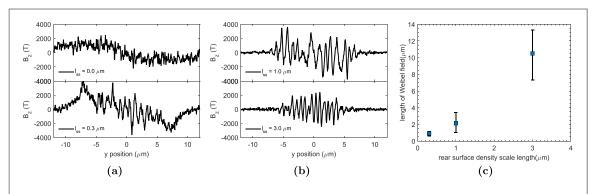
The effect of a range of plasma scale lengths from 0.0  $\mu$ m (plane foil) to 3.0  $\mu$ m are shown in figure 3. Broadly it is observed that in the presence of longer scale length plasmas, magnetic fields build to a similar peak strength as is also shown in figures 4(a) and (b), however the length of the filaments and their lifetime are found to be dependent on the scale length of the plasma, with longer scale length plasmas hosting larger scale and longer lived filamentation of the fields, as shown in figure 3 and quantified in figure 4(c).

The mechanism by which the Weibel instability evolves is consistent with that described by Silva *et al* [23], where the background electron density has a pivotal role. Equation (1) expresses the plasma condition for Weibel instability growth derived by Silva *et al* in terms of the simulation parameters, for a hot electron beam of number density,  $n_b$ , with momenta components,  $\mathbf{p}_{x,y}$ , and Lorentz factor,  $\gamma_b$ , propagating through a background plasma density,  $n_e$ .

$$\frac{n_{\rm b}}{n_{\rm e}} > \gamma_{\rm b} \left(\frac{\mathbf{p}_{y}}{\mathbf{p}_{x}}\right)^{2}.\tag{1}$$

In the case of the 3.0  $\mu$ m rear surface scale length, a clear filamentation of the relativistic electrons is first observed at a time of 650 fs, soon after the peak of the pulse interacts with the target, with peak electron beam densities around the critical density being injected into the target. The electrons have an exponential spectra, but on average are found to have a Lorentz factor in the region of 1.5–2.0, which is consistent with the ponderomotive potential and these electrons have an average momentum ratio,  $\frac{P_y}{P_z}$ , of approximately 0.3. The result is that an Weibel instability can be seeded at a background plasma density of around 7 times the critical density. However, the instability is dynamic, and the generation of magnetic fields at low density quickly leads to the evolution of magnetic fields at higher densities as the incoming beam is pinched by the magnetic fields to higher densities, raising the background plasma density that the condition is satisfied for. Coupled with the energy spread of the beam, and the temporally evolving number density injected, many regions of the plasma can





**Figure 4.** (a), (b): the spatial scale and local values of the magnetic fields at a position of background electron density of 10 times critical density and at a time of 700 fs. (c) The average length of the magnetic field structures in x, measured across the central 2  $\mu$ m in y, are also shown, where error bars give the range in the measured length.

simultaneously satisfy the condition for Weibel instability growth, and this is the basis for the magnetic field channel length.

Locally, the fields reach similar maximum values, around 1.5–3.0 kT, and a lineout across the fields is shown in figures 4(a) and (b) saturating when the field satisfies the gyrotron frequency of the relativistic electrons.

In the vicinity of the instability, the ratio of magnetic pressure to the thermal pressure of the plasma,  $\beta$ , is less than unity. The resultant pressure prevents the magnetic channels from collapsing, and instead forces it to disperse over time.

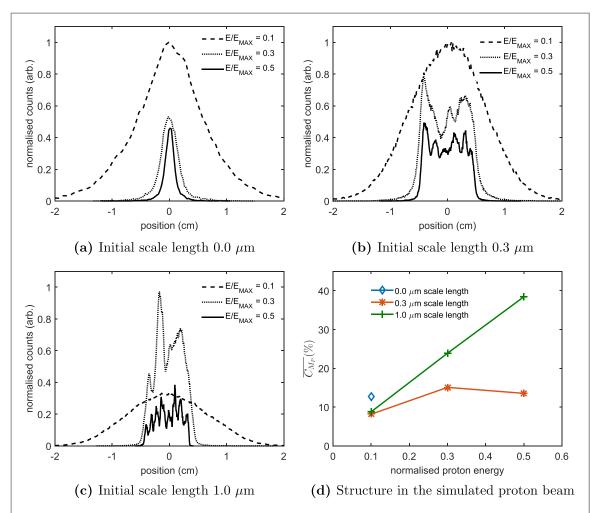


Figure 5. (a)—(c) The proton number as a function of position at the detector plane, as projected from their position and momenta in PIC simulations at a time of 3.8 ps. (d) The structure in the beam as a function of proton energy for the three scale lengths is also presented.

The resultant distribution of the accelerated proton beam in the PIC modelling was calculated by using each particle position and momenta and interpolating its trajectory to a detector plane, 51 mm from the nominal target rear surface. The distributions are shown in figure 5, where the proton energies are normalised to the maximum energy recorded in the simulation of a 1  $\mu$ m scale length plasma.

Qualitatively, the key trends experimentally measured of  $\overline{C_{M_p}}$  as a function of scale length in figure 1(c) and proton energy in figure 2 are reproduced. When no scale length is initially present, a smooth distribution is obtained for three sample energies and the beam is observed to have a larger number of protons at lower energies, as typically observed in TNSA experiments and models.

When a 0.3  $\mu$ m scale length is introduced, significant density perturbations are observed in the higher energy proton beam and is caused by the filamentation of the fast electron beam. However, the lower energy beam is smooth in nature, which is accelerated at late time when the magnetic fields have decayed.

We note that ion acceleration from a non zero density scale length can be complex [24–26, 27], however, in general the time taken for the completion of the ion acceleration becomes longer with increasing density scale length. Here, the increased acceleration time is consistent with the breaking time given by Grismayer and Mora [24]. This forecasts that the acceleration from a 0.3  $\mu$ m plasma scale length will require around 2.5 times longer than a zero density scale length plasma, which itself has a typical acceleration time empirically found to be around 1.3 times the laser pulse duration [28]. This corresponds to a global simulation time of around 1.9 ps in the 0.3  $\mu$ m case, while the strong filamentary structure in the magnetic field is observed to have faded at an earlier time of around 1.0 ps, and so the lower energy protons are not affected by these structures as they are accelerated toward the end of this time window.

For a longer scale length of 1.0  $\mu$ m, a similar degree of filamentation in the proton beam is observed at intermediate and low energy, while the high energy beam is filamented to a greater degree than for the shorter plasma scale length. In this instance, the proton acceleration is expected to take just over six times as long as the zero density scale length case, corresponding to a time of around 4.4 ps in this simulation. This longer

acceleration time explains the deviation of the number of accelerated protons from the exponential distribution expected, as the ion acceleration process has not quite finished.

In the 3.0  $\mu$ m scale length case, the acceleration time is estimated to be around 12 ps, and the ion energies are observed to be much lower than in other simulations after 3.8 ps. For this reason a detailed comparison at similar energies is not possible, and is therefore not presented.

To quantify the degree of structure in the simulated proton beam, a similar method to that used on the RCF stacks was used, but as the projected proton beam profile is two dimensional, no radial averaging takes place. The averaged coefficient of variation across the central 20, 10 and 8 mm of the beam is shown for normalised proton energies of 0.1, 0.3 and 0.5 respectively in figure 5(d).

For all scale lengths investigated, an approximately Gaussian beam is obtained for the lowest energy component, and this yields a small averaged coefficient of variation,  $\overline{C_{M_p}}$ , at low energies in good agreement with the values obtained in figure 2. The higher energy component of the zero scale length component also forms a smooth beam, but with a profile not in agreement with the experimental results, which are better approximated by a parabolic or top-hat envelope, which is observed for the other simulated profiles. In absence of a similar global profile, the evaluation of  $\overline{C_{M_p}}$ , reflects changes in the global structure of the beam rather than the substructure of the beam that we aim to examine, and so results for the higher energy components are omitted.

Qualitatively, good agreement in the trends observed in the experimental data are reproduced, where a 0.3  $\mu$ m scale length plasma produces a beam with an approximately constant  $\overline{C_{M_p}}$  as observed in the 10 ps delay experimental data set, and the 1.0  $\mu$ m scale length plasma produces a beam with  $\overline{C_{M_p}}$  steadily increasing as a function of energy, as observed in the 100 ps delay experimental data set.

# 3.2. Three dimensional particle in cell

To investigate the stability of this mechanism in three dimensions, the EPOCH code was again used, with a box of dimensions  $13(x) \mu m \times 10(y) \mu m \times 10(z) \mu m$ , with 20(x) and 40(y,z) cells per micron, where the plasma had the same densities and composition as in the 2D simulations, with 25 particles per cell and a laser spot size of 1.8  $\mu m$ .

When a 1  $\mu$ m scale length was investigated, the same magnetic field structures observed in 2D modelling are found in the xy midplane, shown in figure 6(a). The structure of these fields can be examined in the acceleration (yz) plane and this is shown in figure 6(b). Once again, as the plasma pressure mainly prevents the collapse of electrons into the magnetic field structure, the electrons form circular structures, with voids and higher density regions. Similar structures to this in the electron distribution have been shown to lead to filamentation of the accelerated proton beam [6]. While modelling the ion acceleration until its conclusion would be computationally expensive for our purposes, the beam obtained at the end of the simulation, shown in figure 6(d), shows some degree of filamentation similar to that observed experimentally, with an angular web like structure in the proton number density emerging in the density map, which are qualitatively similar to those observed in experiment.

#### 4. Discussion

## 4.1. General discussion

Evidence has been presented for the case that a Weibel instability in a rear surface plasma density gradient is the source of a filamented proton beam observed in experiment.

Numerical modelling shows that increasing the rear surface plasma scale length changes the Weibel instability evolution in two important ways. Firstly, it is shown that a scale length is required for the instability to be seeded, and that the size of the scale length determines the filament length in the rear surface plasma. The filamentation of fast electrons via the Weibel instability results in a filamentation of the proton beam during the ion acceleration phase, which is both experimentally measured and predicted by PIC simulations. By controlling the inter pulse time delay in the experiment, we therefore control the plasma scale length into which the hot electrons are injected, and whether the Weibel instability is present or not.

Secondly, the length of the filaments is the basis for the stability of the magnetic field and for its lifetime, leading to longer density scale lengths hosting magnetic field filaments for longer time. The result of this is that the fast electrons remain filamented for longer times in the acceleration phase, and that the degree of structure mapped to the proton beam varies as a function of proton energy. This is once again measured clearly in the experimental results, where longer delays between the pulses led to a higher degree of structure being mapped onto lower energy protons, as shown in figure 2 and qualitatively reproduced from 2D PIC simulations in figure 5(d).

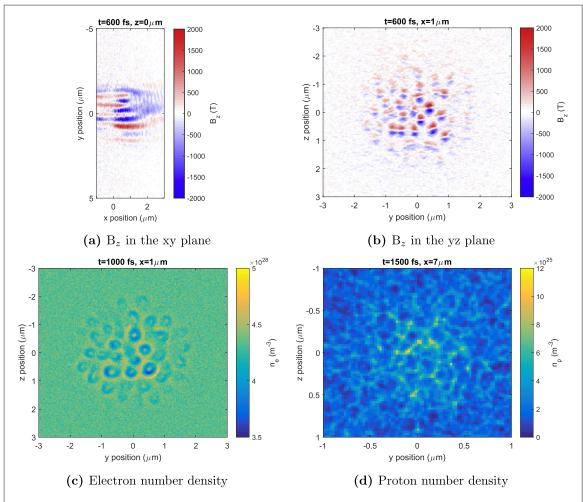


Figure 6. The magnetic field evolution in 3D PIC simulations and its influence on the electron and proton density distributions in the acceleration plane.

No experimental measurements were made of the scale lengths, however the development of a scale length of 0.3  $\mu$ m at times greater than 3 ps is consistent with that expected for a self similar expansion of  $\approx$ 300–400 eV fully ionised carbon, which could be considered a conservative estimate of the bulk target conditions initiated by the controlled prepulse, as will be discussed in section 4.2.

A gold target would be unlikely to be fully ionised, and therefore a slower expansion than the carbon target could be expected, with shorter density scale lengths for the same plasma temperature, and figure 1(c) appears to agree with this hypothesis, showing a slower onset of structure in the beam with inter pulse time delay.

This effect could be potentially important for high energy multi picosecond systems, like those envisaged for fast ignition experiments, where uniformly heating a hot spot is of key importance and this work demonstrates that on these time scales, ion beam and electron beam uniformity is affected by the existence of such scale lengths.

In general, ion based fast ignition experiments could conceivably involve some degree of plasma scale length on the rear of the ignitor beam target, preheated by x-rays produced from the primary fuel compression stage, making this work relevant not only to 10 s of picosecond ignitor beams.

#### 4.2. Discussion of the experimental scale length estimation

A precise estimation of the target rear expansion velocity, and hence the target rear scale length evolution, is not trivial, as to our knowledge there an no dedicated studies in the literature on the scaling of this parameter when only the front side of the target is irradiated, although passing references to the phenomenon have been made [29].

However, to relate the experimental data to the scale lengths in the numerical investigation, we discuss this in further detail, and draw from various sources. Although the prepulse intensity is a little higher than could be considered for a purely collisional heating model, the scaling of Gibbon's analytic expression [30] of the collisional skin heating model by Rozmus and Tikhonchuk [31] yields a plasma temperature in the region of 1.8 keV, whilst a 2D collisional PIC model of parameters very close to that reported for the prepulse in this work,

by Sentoku *et al* [32] predict a front surface temperature of approximately 3–4 keV. These compare with the experimentally measured peak temperatures of 1.5 keV in the laser spot irradiated area on the front surface, with peak intensities slightly higher than those reported here at  $2 \times 10^{18}$  W cm<sup>-2</sup> [33].

In a separate study with hybrid PIC modelling [6], the bulk electron temperature was found to fall to around 10% of the front surface value over the thickness of the carbon foils used in this study, when heated by an injected electron beam. Combining this with the discussed range of possible front surface plasma temperatures would imply a rear surface temperature range of 150–400 eV.

Further than this, in a recently developed technique by Bocoum *et al* [34], it is found that a hundreds of eV temperature plasma expands with a constant sound speed at the critical surface, for measured times of up to 30 ps.

For fully ionised carbon, this temperature range would infer scale lengths of 0.25–0.42  $\mu$ m at a time of 3 ps, growing linearly up until at least a time of 30 ps, where they would reach 2.5–4.2  $\mu$ m.

It is reasonable to expect that the front surface preplasma evolves due to a self similar expansion driven by the hotter plasma temperature there, and in the case of the work on the 45  $\mu$ m carbon foils have an expansion of around three times faster than at the back surface, owing to the  $\sqrt{T_{\rm e}}$  scaling. This would infer maximum front surface scale lengths in the region of 1.3–13  $\mu$ m at 3–30 ps, and potentially up to 40  $\mu$ m scale lengths at times of up to 100 ps.

#### 4.3. Discussion of possible front surface sources of filamentation

We have shown that micron scale lengths are capable of hosting large Weibel instabilities, and this instability could theoretically form in the front surface scale lengths of the interaction, a phenomenon not included in the numerical modelling here. Whilst the condition for Weibel instability growth will be satisfied in the lower density preplasma, at sufficiently large background plasma densities, especially in the experimental case, where the background plasma densities can reach up to 500 times the critical density, the condition is not satisfied, leading to ballistic electron transport through the target. To determine whether a front surface seeded structure could be ballistically transported through the bulk target, we consider a filamented electron beam with an average periodicity of  $\bar{a}$  and momenta  $\mathbf{p}_{x,y}$ . The distance they must travel in the x direction such that the neighbouring filaments overlap, d, is:

$$d = \bar{a} \ (\frac{\mathbf{P}_x}{\mathbf{P}_y}). \tag{2}$$

If the target thickness is greater than d, it can be considered too thick for the filamentary structure to be transported through the bulk. From 2D PIC modelling on the target rear, these quantities could be expected to take of value of  $\bar{a}=250\,$  nm and  $\mathbf{p}_y/\mathbf{p}_x=1/3$ , and this gives a threshold thickness of  $\approx\!800\,$  nm which is very much smaller than the typical target thickness used in this study.

Scale lengths of  $10-40~\mu m$ , predicted to be present on the target front surface for the longest time delays have been subject to experimental and numerical investigation [12–16], where it has been found that these scale lengths are not long enough for self focussing effects to lead to laser beam filamentation, allowing us to conclude that this effect has no role in any front surface seeding of structure in the electron beam.

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