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Energy coupling in short pulse laser solid interactions and its impact for space debris removal

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Significant advances have been made over the last decade to improve the performance, efficiency and contrast of high peak and average power laser systems, driven by their use in a wide variety of fields, from the industrial to the scientific. As the contrast of the lasers has improved, interactions with contrasts of 10^{-12} are now routinely undertaken. At such high contrasts, there is negligible pre-plasma formation and the ionized surface layer created by sub-ps duration pulses typically forms a highly reflective “plasma mirror” capable of reflecting between 70–90% of the incident energy. Although such interactions are of significant interest for applications such as harmonic source production and to enable the underlying physics to be studied, their low absorption can limit their usefulness for applications such as space debris removal.

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1. Introduction

The growing demand for utilization of satellites by governments, companies and individuals is increasingly threatened by collisions with debris in low earth orbit (< 2000 km). The potential for a self-sustaining cascade of collisions known as the Kessler syndrome [1] is increasing as the number of satellites and pieces of space debris has risen over the last four decades. There is a growing realization [2-3] amongst the community capable of orbital launches, that all future satellites must have adequate maneuvering capability so that they can be placed in a mid-term (25 year) decay orbit after their intended operational life, or safely de-orbited [4] by other techniques [5].

A laser based system for de-orbiting either the space debris or satellites after the end of their life cycle was first proposed [6] almost 20 years ago. Although the recoil from photon reflection/absorption could be used to de-orbit an object, the momentum imparted by laser ablated material is typically ~10^{4} times larger for a given quantity of photons, if the power density is sufficient to cause ablation. Therefore, most designs have considered a high power laser system either terrestrially based or in orbit which could be repeatedly focused onto an object causing material ablation and an orbital velocity change. Recent advances in laser technology over the last two decades have significantly increased the repetition rate and scope of available optical parameters which could be delivered. In the field of high intensity systems, the peak optical powers have typically increased from TW to PW’s, with pulse lengths reducing from ns to ~100’s fs. The laser pumping technology is also advancing with the introduction of Diode Pumped Solid State Lasers [7]. Here, the emission wavelength of the pumping source can be matched to the lasing material absorption to ensure efficient coupling and electrical to lasing efficiencies of 10-20% are now deliverable.

In this paper, the role of the plasma density scale length present during the interaction of short sub ps pulses on laser energy absorption will be explored and the consequences for orbital debris removal discussed.

2. Laser debris removal system requirements

When an intense laser pulse illuminates a surface at a fluence \( \Phi (J/m^2) \), it is possible to cause ablation if the fluence exceeds the threshold for that material \( \Phi_{t} \). Optical ablation has been studied by many groups [8] and it was generally found that the fluence required for plasma ignition \( \Phi_{t} \sim 5 \times 10^{6} \text{eV}/\text{cm}^2 \), where \( \tau \) is the duration of the illuminating pulse \( \Phi > \tau \). The damage threshold measured at wavelength \( \lambda = 1.055 \times 10^{-5} \text{m} \) (the wavelength of Nd-doped glass lasers) in the regime 100 ns > \( \tau > 1 \) ps for a number of materials is shown in Figure 1. It can be seen, that when \( \tau > 100 \) ps the damage threshold is a vertical line on
Figure 1 for a given material. In this regime of pulse lengths, damage is considered to be caused by energy absorption and transfer to the material's lattice. As thermal conduction dominates the heat flow, a $r^{0.5}$ dependence on $\Phi(t)$ is to be expected [10]. For pulses of a few ps and less, the energy is deposited into a surface layer for an absorbing metal and heating above the boiling point is sufficiently rapid that negligible heat conduction takes place. Hence, the line on Figure 1 becomes nearly horizontal for pulses $<1$ ps. Intermediate between 1 and 100 ps, there is a changeover region for the heat flow and this can be observed as a smooth transition, the exact location of which is material dependent.

As well as being above the ablation threshold, the efficiency with which a given laser pulse generates pressure must be understood before a choice of suitable design parameters for an orbital debris removal system can be selected. We can use the momentum coupling parameter $C_m$ which is the pressure $p$ exerted on the target divided by the Intensity $I$ (Wm$^{-2}$) [3]. Phipps [11] has correlated data from a number of studies at various wavelengths and pulse durations and the trend line for $C_m$ is plotted against $I/r^{0.5}$ as shown in Figure 1. If the product of $I/r^{0.5}$ is too low ($<100$ Jm$^{-2}$), thermal conduction works to limit the temperature rise and no ablation occurs. As $I/r^{0.5}$ is increased, significant surface heating occurs and the material may vaporize if $\Phi>\Phi_0(t)$ causing $C_m$ to rapidly rise to a maximum. Initially, the laser can still penetrate through the transparent vapor and deposit energy directly into the original surface. However, there is a rapid transition to absorbing plasma and the laser now deposits energy above the surface, imparting force on the target through the plasma pressure. As the temperature of the plasma increases, a greater fraction of the incident energy is used to ionize the plasma and more and more energy is lost to escaping radiation, hence, $C_m$ slowly decreases. For ultra-high intensities ($I/r^{0.5} > 10^9$) the light pressure can dominate over the plasma pressure and push the absorbing surface back towards solid densities. However, at such intensities most of the energy is transported away from the focus in relativistic electrons [12], which can deposit their energy far into the target and $C_m$ is low.

The extreme temperatures and gradients associated with short pulse illumination can also cause the expulsion of a wide variety of states of matter from ions [13], atoms, clusters, nanoparticles [14] and droplets. Amoruso et al. [14] found that at intensities a few times above threshold, three main components to the expanding plasma were identified: a low number of fast ions at $<10^3$ ms$^{-1}$; a much larger neutral and ion population at $\sim 10^4$ ms$^{-1}$; and a much slower component at $10^6$ ms$^{-1}$ containing 10's nm sized nanoparticles. Using a simple model based on assuming that the dominant absorption mechanism was due to avalanche ionization, good estimates for the ablation thresholds were obtained in the sub-ps duration interaction regime. However, it is only the trailing edge of the pulse that is strongly absorbed and Amoruso also noted that strong reflection could occur.

3. Plasmas as reflectors

When studying the absorption of short pulse (~ps) intense lasers with solid matter, it has long been realized that the presence of a pre-plasma can significantly change the interaction physics [15-16]. Although the absorbed energy $E_a$ can be determined from measurements made using an integrating sphere, in which all the scattered, transmitted [17] or reflected laser energy is collected and subtracted from the incident energy $E$, this is intrusive as it requires a large solid angle, blocking any other diagnostic measurements. A more common technique implemented on many experiments is to measure the specularly reflected laser energy $E_{spec}$ and assume $E_a=E-E_{spec}$ in the absence of significant scatter or transmission. For medium-to-high contrast laser systems, where the intensity contrast is typically $>10^6$ [18] and can be up to $10^{12}$ a few ps before the arrival of the main pulse, it has been found that the ionized surface can act as a very efficient [19] almost optically flat “plasma mirror” during the interaction [20]. The specular reflectivity of fused silica measured by Ziener [19] for 90 and 500 fs duration pulses at 800 nm is shown in Figure 2. As expected, the switchon in reflectivity for the shorter pulse 90 fs data is at a slightly higher of $I/r^{0.5}$ than the 500 fs data due to the scaling of $\Phi(t)$ for shorter pulses [8]. As the intensity is increased above threshold, the reflectivity increases and peaks at $\sim 70%$ for $I/r^{0.5}>10^4$ Jm$^{-2}$ and then drops rapidly for $I/r^{0.5}>10^5$ Jm$^{-2}$. This fall-off in reflectivity is associated with the plasma being sufficiently hot that it can significantly expand hydro-dynamically during the period of illumination. Also shown on Figure 2 is the optical transmission...
of fused silica measured for 1.5 and 3 mm samples. The transmitted energy through the 3 mm sample shows a lower threshold value and lower transmission (once above threshold). This effect is associated with self-focusing as the beam passes through the sample which can cause filamentation down to diameters on the order of $\lambda$, and locally raises the fluence above $\Phi_e$, causing energy deposition within the bulk of the media.

Experiments have also been carried out where the level of laser energy preceding the main 50 fs pulse has been varied and the specular reflectivity measured [21]. Although these measurements were carried out for rather high values of $I_0^{0.5} > 10^8$ they clearly demonstrate, that a larger pre-plasma present before the main pulse can significantly reduce the plasma reflectivity and hence potentially increase absorption. However, at values of $I_0^{0.5}$ closer to the peak for momentum coupling, the energy required to generate a pre-plasma with large scale length could represent a significant fraction of the energy contained in the main peak. Work has been carried out using such fs/ns pulse combinations [22, 23] for ablation studies and it was generally found to be advantageous. However, specific experiments would need to be undertaken to find if a more efficient regime for impulse generation can be found.

The reflectivity for a metal, in this case gold, measured for a 0.5 ps pulse of wavelength 1.054x10$^{-6}$ m from below the ablation threshold to $I_0^{0.5} = 10^8$ Jm$^{-1}s^{-0.5}$ for S polarized laser light is shown in Figure 3. This curve shows very high reflectivity >90% for low values of $I_0^{0.5} < 10^6$ Jm$^{-1}s^{-0.5}$ leaving the maximum energy deposited into the plasma to be < 10%. Therefore, there will be very low coupling efficiency into a plasma. This should be contrasted against the case of using long (~ns) pulses where the absorption is dominated by Inverse Bremsstrahlung heating and absorption fractions into the plasma of >25% are common. The reflectivity for the short pulse case only drops significantly for very high values of $I_0^{0.5} > 10^9$ Jm$^{-1}s^{-0.5}$. This would again be unsuitable for generating efficient momentum.

4. Conclusion

Short duration illumination (<1ps) where there is limited heat conduction into the target compared to the ns regime was considered to be advantageous for both laser propulsion and debris removal [11] in the mid 1990s. Recent measurements demonstrate that, for pre-pulse-free laser interactions with solid matter above the ablation threshold, the majority of energy is reflected rather than being absorbed. This is found to be generic across many systems where the pulse length is in the 50 fs – ~ps regime and the intensities are within a few orders of magnitude above the ablation threshold $\Phi_e$. For intensities many orders of magnitude above $\Phi_e$, different absorption mechanisms can dominate. However, at $I_0^{0.5} > 10^9$, for sub ps interactions, the electron and ion temperatures become significant and only a small fraction of the energy goes into producing momentum. New studies measuring the momentum coupling using ns, low contrast or combined ps/ns laser systems will help determine the optimum interaction conditions for orbital debris removal.

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